

# Simulation of the generation of the characteristic X-ray emission by hot electrons in a foil

O.F. Kostenko, N.E. Andreev

**Abstract.** We have developed a model to calculate the yield of the characteristic X-ray radiation from a foil, taking into account the dependence of the average energy and the number of hot electrons on the intensity of the laser pulse, the self-absorption of X-rays and the effect of refluxing of hot electrons. The yield of  $K_\alpha$  radiation from a silver foil is optimised at relativistic intensities. A method is proposed for diagnosing the effect of electron refluxing, which greatly increases the yield of  $K_\alpha$  radiation.

**Keywords:** hot electrons, refluxing, generation and absorption of  $K_\alpha$  radiation, foil.

## 1. Introduction

Hot electrons produced by irradiating a solid target by a relativistically intense laser pulse [1] penetrate into the target and ionise the K-shell of atoms, which leads to the emission of characteristic X-rays on  $2p-1s$  transitions in excited atoms. The generated short-pulse  $K_\alpha$  radiation can be used, for example, to diagnose a dense plasma (see reviews [2, 3]). Of importance here is the problem of increasing the number of generated photons and their energy [2]. The  $K_\alpha$  yield can be increased by increasing laser pulse energy and efficiency of its conversion to the energy of this radiation. At relativistic intensities, the efficiency, achieved through the use of isolated thin foils with limited transverse dimensions, in which hot electrons are trapped and refluxing, is much greater than in the case of foils placed on a massive conductor, which due to the return current of cold electrons compensates for the charge of fast electrons leaving the foil [3–5]. In these papers the results of measurements of the coefficient of the laser pulse energy conversion into the  $K_\alpha$  radiation energy were compared with the results of model calculations, taking into account the loss of energy of the hot electrons in the foil, the cross section of the K-shell ionisation by an electron impact and the probability of  $K_\alpha$  radiation de-excitation by excited atoms. The energy distribution of the electrons impinging onto the foil was determined by measuring the bremsstrahlung spectrum and by using calculations by the particle-in-cell (PIC) method [3] or was considered exponential with the temperature defined by the ponderomotive potential [3–5]. In the latter case it was assumed that the electron temperature  $T_h$

depends on the peak intensity, and the conversion efficiency of the laser energy into the energy of hot electrons  $\eta$  is constant over a wide range of intensities – from sub-relativistic to ultra-relativistic.  $K_\alpha$  radiation absorption in thin foils was assessed using a constant factor.

In this paper, we construct a model that takes into account the  $K_\alpha$  radiation absorption in the foils of arbitrary thickness with and without electron refluxing and the dependences of  $\eta$  and  $T_h$  on the laser pulse intensity. We have obtained the dependences of the anisotropic photon yield from the front and rear sides of the foil (per unit laser energy and per unit solid angle) on the foil thickness and the peak intensity. The calculations have been performed for a silver foil (the photon energy of  $K_\alpha$  radiation is 22.1 keV) at intensities corresponding to the experiment conducted on the PHELIX laser facility [3].

## 2. Description of the model

An electron with energy  $E_0$ , incident on a foil of thickness  $d$  (perpendicular to its surface), generates

$$dn_K = \omega_K p_\alpha n_a \sigma_K(E) dx$$

$K_\alpha$  photons on a path  $dx$  at a distance  $x$  from its surface, where  $\sigma_K(E)$  is the impact ionisation cross section of the K-shell by an electron with energy  $E(E_0, x)$ ;  $n_a$  is the concentration of atoms;  $\omega_K$  is the probability of radiative de-excitation; and  $p_\alpha$  is the probability of the  $K_\alpha$  fluorescence yield. In this case, from front side of the foil at an angle  $\alpha_0$  to the normal

$$dn_{em} = dn_K \frac{d\Omega}{4\pi} \exp\left(-\frac{x}{l_a \cos \alpha_0}\right)$$

photons are emitted into a solid angle  $d\Omega$ , where  $l_a$  is the absorption length. The electron energy loss is described by the function  $S_p(E)$ :

$$\frac{dE}{dx} = -S_p(E).$$

The electron mean free path in a solid, where its energy decreases from  $E_0$  to the ionisation potential of the K-shell  $E_K$ , is

$$l_e(E_0) = \int_{E_K}^{E_0} \frac{dE}{S_p(E)},$$

whence  $x = l_e(E_0) - l_e(E)$ .

If  $l_e(E_0) \ll d$ , then the total number of photons emitted by the electron per unit solid angle in the direction determined by the angle  $\alpha_0$  is

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$$\frac{dN_{\text{em}}^f(E_0)}{d\Omega} = \frac{\omega_K p_\alpha n_a}{4\pi} \int_{E_K}^{E_0} \frac{\sigma_K(E)}{S_p(E)} dE$$

$$\times \exp\left[-\frac{l_c(E_0) - l_c(E)}{l_a \cos \alpha_0}\right]. \quad (1)$$

Therefore, the number of the photons emitted from the back side of the foil per unit solid angle at an angle  $\beta_0$  to the normal is equal to

$$\frac{dN_{\text{em}}^b(E_0)}{d\Omega} = \frac{\omega_K p_\alpha n_a}{4\pi} \int_{E_K}^{E_0} \frac{\sigma_K(E)}{S_p(E)} dE$$

$$\times \exp\left\{-\frac{d - [l_c(E_0) - l_c(E)]}{l_a \cos \beta_0}\right\}. \quad (2)$$

If the electron, after reaching the back side of the foil at  $l_c(E_0) > d$ , leaves it, then in Eqns (1) and (2) the lower limit of integration  $E_K$  is replaced by the energy  $E_i$ , which is found from the solution of the equation  $l_c(E_0) - l_c(E_i) = d$ . If the electron refluxes in the foil at  $l_c(E_0) > d$ , then

$$\frac{dN_{\text{em}}^f(E_0)}{d\Omega} = \frac{\omega_K p_\alpha n_a}{4\pi} \sum_{i=1}^{N+1} \int_{E_i}^{E_{i-1}} \frac{\sigma_K(E)}{S_p(E)} dE$$

$$\times \exp\left\{\frac{(2p_i - 1)[l_c(E) - l_c(E_{i-1})] - (1 - p_i)d}{l_a \cos \alpha_0}\right\}, \quad (3)$$

$$\frac{dN_{\text{em}}^b(E_0)}{d\Omega} = \frac{\omega_K p_\alpha n_a}{4\pi} \sum_{i=1}^{N+1} \int_{E_i}^{E_{i-1}} \frac{\sigma_K(E)}{S_p(E)} dE$$

$$\times \exp\left\{\frac{(1 - 2p_i)[l_c(E) - l_c(E_{i-1})] - p_i d}{l_a \cos \beta_0}\right\}, \quad (4)$$

where  $N = [l_c(E_0)/d]$ ;  $p_i = i - 2[i/2]$ , i.e.,  $p_i$  is equal to unity if  $i$  is an odd number, and to zero if  $i$  is an even number. When  $i \in 1, \dots, N$ , the energy  $E_i$  is found by solving the equation  $l_c(E_0) - l_c(E_i) = id$ , and  $E_{N+1} = E_K$ .

The spectrum of the hot electrons incident on the foil was described by an exponential energy distribution

$$dn_e = \frac{dN_e}{T_h(I_L)} \exp\left[-\frac{E_0}{T_h(I_L)}\right] \quad (5)$$

with an average energy determined by the ponderomotive potential [1]

$$T_h(I_L) = mc^2 \{[1 + a_L(I_L)^2]^{1/2} - 1\},$$

where  $a_L(I_L) = 0.855 (I_L/10^{18} \text{ W cm}^{-2})^{1/2} (\lambda/1 \mu\text{m})$  is the normalised amplitude of the laser field;  $I_L(r, t)$  is the intensity of the laser pulse in the focal spot; and  $\lambda$  is the wavelength. The number of electrons  $dN_e$  was found from the law of conservation of energy

$$dN_e T_h(I_L) = \eta(I_L) dE_L, \quad (6)$$

where  $\eta(I_L)$  is the conversion efficiency of the laser energy

$$dE_L = I_L(r, t) 2\pi r dr dt \quad (7)$$

into the energy of hot electrons.

It follows from formulas (5)–(7) that the laser pulse generates

$$\frac{dN_K}{d\Omega} = 2\pi \int_0^\infty r dr \int_{-\infty}^\infty dt \frac{\eta(I_L) I_L(r, t)}{T_h^2(I_L)} \int_{E_K}^\infty dE_0$$

$$\times \exp\left[-\frac{E_0}{T_h(I_L)}\right] \frac{dN_{\text{em}}(E_0)}{d\Omega} \quad (8)$$

photons per unit solid angle, where  $dN_{\text{em}}(E_0)/d\Omega$  is defined by expressions (1)–(4), as appropriate. In the case of the Gaussian intensity distribution

$$I_L(v) = I_0 e^{-v}, \quad v = \frac{r^2}{r_0^2} + \frac{t^2}{t_0^2}$$

formula (8) takes the form

$$\frac{dN_K}{d\Omega} = 2\pi r_0^2 t_0 I_0 \int_0^\infty \sqrt{v} dv \frac{\eta(v) e^{-v}}{T_h^2(v)}$$

$$\times \int_{E_K}^\infty dE_0 \exp\left[-\frac{E_0}{T_h(v)}\right] \frac{dN_{\text{em}}(E_0)}{d\Omega}.$$

It follows from (7), the laser pulse energy is  $E_L = \pi^{3/2} r_0^2 t_0 I_0$ ; therefore,

$$N_{K\alpha} = \frac{1}{E_L} \frac{dN_K}{d\Omega} = \frac{2}{\sqrt{\pi}} \int_0^\infty \sqrt{v} dv \frac{\eta(v) e^{-v}}{T_h^2(v)}$$

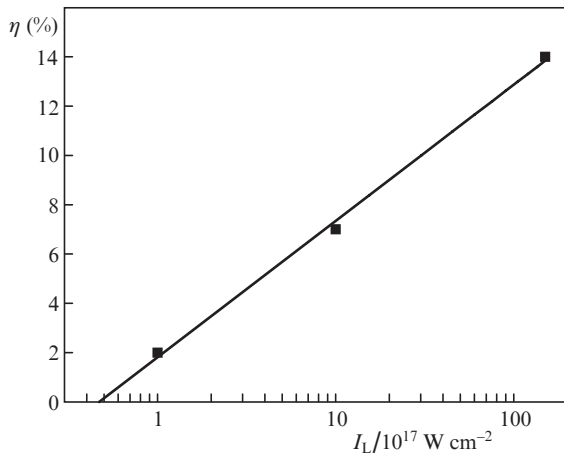
$$\times \int_{E_K}^\infty dE_0 \exp\left[-\frac{E_0}{T_h(v)}\right] \frac{dN_{\text{em}}(E_0)}{d\Omega} \quad (9)$$

photons per unit laser energy are emitted in a given direction per unit solid angle.

### 3. Simulation results

Calculations by formula (9) were performed at intensities  $I_0 = 5 \times 10^{17} - 1.5 \times 10^{19} \text{ W cm}^{-2}$  ( $\lambda = 1.053 \mu\text{m}$ ), which correspond to experiments carried out on the PHELIX laser facility with a pulse duration of  $\sim 700$  fs [3]. In accordance with paper [6], in the case of normal incidence of a high intensity laser pulse ( $I_L \lambda^2 \approx 10^{17} \text{ W cm}^{-2} \mu\text{m}^2$  and higher) the generation of hot electrons is due to the oscillatory component of the ponderomotive force directed along the gradient of concentration of thermal electrons. The comparison of the dependences of hot electron temperature on  $I_L \lambda^2$  with the results of measurements performed upon interaction of the different femtosecond laser pulses with solid targets, shows [7] that at such intensities there occurs a transition to ponderomotive scaling [1]. Laser energy conversion coefficient  $\eta(I_L)$  was determined by fitting the results of PIC calculations (Fig. 1), which were carried out for a 700-fs laser pulse incident at a small angle. The integration interval in the variable  $v$  is limited from above by  $v_{\text{max}} = \ln(I_0/I_L^*)$ , which corresponds to the condition  $\eta(I_L) = 0$  at  $I_L \leq I_L^* \approx 5 \times 10^{16} \text{ W cm}^{-2}$ . The upper limit of integration over energy  $E_0$  is chosen equal to  $9T_h(v)$ , taking into account the results of measurements and PIC simulations [3].

Energy loss of electrons in silver  $S_p(E)$  was calculated using the ESTAR database [8]. The cross section of the K-shell ionisation,  $\sigma_K(E)$ , is determined according to the analytical expression [9], in which the corresponding factor [10, 11] takes into account the growth of the cross section with increasing energy of the relativistic electrons. In the calcula-

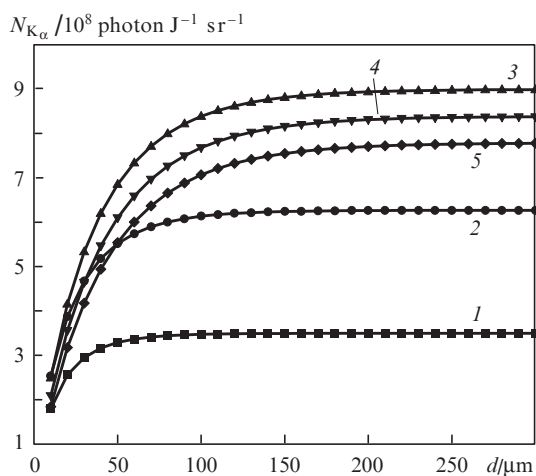


**Figure 1.** Dependence of the laser energy conversion efficiency into the energy of hot electrons on the intensity: points – calculation by the PIC method [3], solid line – approximation.

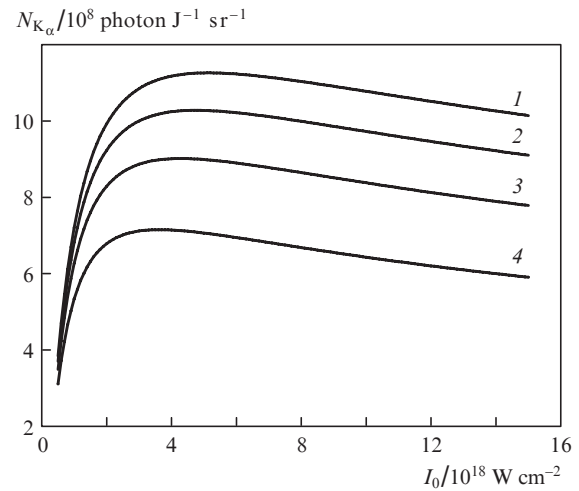
tions the following parameters were used: the probabilities  $\omega_K = 0.831$  [12] and  $p_\alpha = 0.826$  [13], the concentration of atoms  $n_a = 5.86 \times 10^{22} \text{ cm}^{-3}$ , the ionisation potential  $E_K = 25.516 \text{ keV}$  and the absorption length  $l_a = 74 \mu\text{m}$  [14] of  $K_\alpha$  radiation with the weighted mean energy of 22.1 keV (defined according to [13, 15]), which correspond to silver under normal conditions.

The  $K_\alpha$  yield from the front side of the foil without electron refluxing increases with increasing foil thickness and tends to a constant value (Fig. 2). The maximum  $K_\alpha$  yield at an observation angle of  $45^\circ$  is reached at an intensity of  $4 \times 10^{18} \text{ W cm}^{-2}$ . When reducing observation angle the maximum yield increases slightly (Fig. 3). The  $K_\alpha$  yield from the rear side of the foil reaches a maximum at some thickness, which increases with increasing intensity (Fig. 4). At optimal thicknesses and the same observation angle, the  $K_\alpha$  yield from the front side of the foil is higher than that from the rear side (Fig. 5).

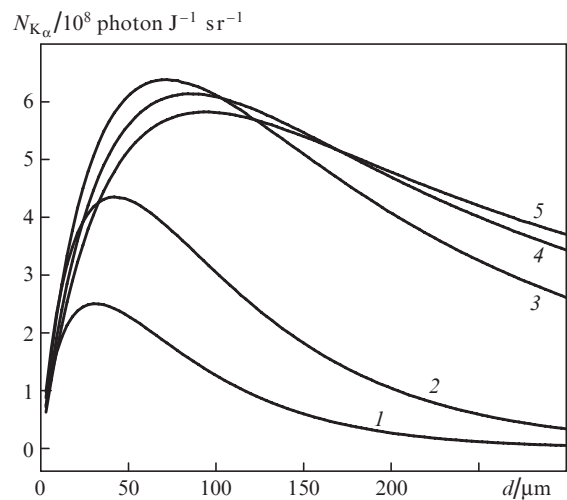
If electrons are refluxed in the foil, the  $K_\alpha$  yield from the front side of the foil increases with increasing intensity and



**Figure 2.**  $K_\alpha$  yield from the front side of the foil without electron refluxing as a function of the foil thickness at peak intensities  $I_0 = 5 \times 10^{17}$  (1),  $10^{18}$  (2),  $5 \times 10^{18}$  (3),  $10^{19}$  (4), and  $1.5 \times 10^{19} \text{ W cm}^{-2}$  (5) and an observation angle  $\alpha_0 = 45^\circ$ .



**Figure 3.**  $K_\alpha$  yield from the front side of the 500- $\mu\text{m}$ -thick foil without electron refluxing as a function of the peak intensity at observation angles  $\alpha_0 = 0$  (1),  $30^\circ$  (2),  $45^\circ$  (3) and  $60^\circ$  (4).



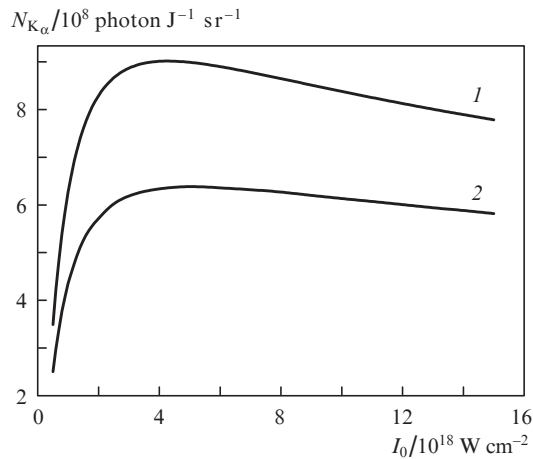
**Figure 4.**  $K_\alpha$  yield from the rear side of the foil without electron refluxing as a function of the foil thickness at peak intensities  $I_0 = 5 \times 10^{17}$  (1),  $10^{18}$  (2),  $5 \times 10^{18}$  (3),  $10^{19}$  (4), and  $1.5 \times 10^{19} \text{ W cm}^{-2}$  (5) and an observation angle  $\beta_0 = 45^\circ$ .

decreasing its thickness (Fig. 6). At  $I_0 \geq 10^{18} \text{ W cm}^{-2}$  and the same observation angles the  $K_\alpha$  yield from the rear side of the foil is less than the yield from the front side with thicknesses greater than the effective absorption length  $l_a \cos \alpha_0$  (Fig. 7).

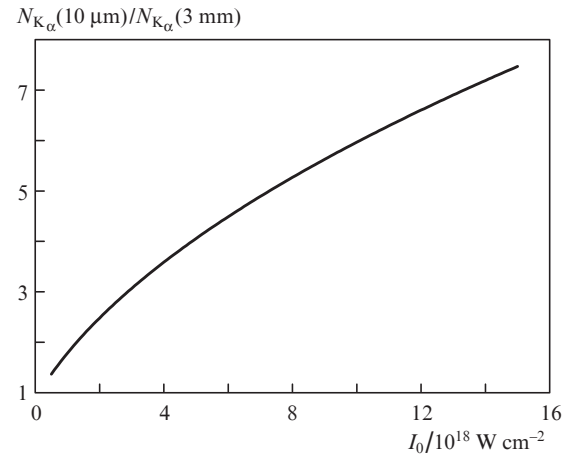
The relative yield of  $K_\alpha$  radiation can be measured with greater accuracy than the absolute yield. The ratio of the  $K_\alpha$  yields from the front side of a thin and thick foils increases with increasing intensity for the target with the electron refluxing taken into account (Fig. 8) and decreases for the target in which refluxing is insignificant (Fig. 9). This feature can be used to study the effect of electron refluxing, which greatly increases the  $K_\alpha$  yield (see Figs 5 and 6).

#### 4. Conclusions

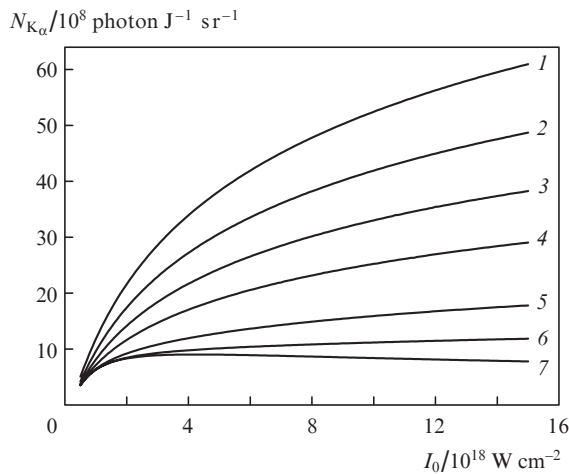
Thus, we have constructed a model of generation of  $K_\alpha$  radiation by hot electrons in a foil, which takes into account the spatial dependence of the probabilities of photon emission



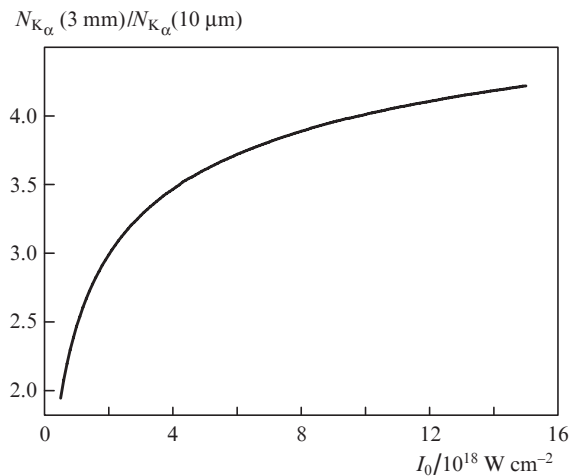
**Figure 5.**  $K_\alpha$  yield from the front (1) and rear (2) sides of the foil without electron refluxing as a function of the peak intensity at the corresponding optimal thicknesses and observation angles  $\alpha_0 = \beta_0 = 45^\circ$ .



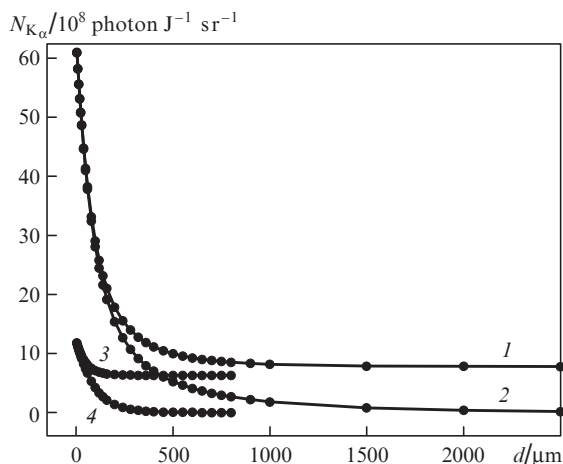
**Figure 8.** Ratio of the  $K_\alpha$  yields from the front side of a thin (10  $\mu\text{m}$ ) and thick (3 mm) foils with electron refluxing as a function of the peak intensity at  $\alpha_0 = 45^\circ$ .



**Figure 6.**  $K_\alpha$  yield from the front side of the foil as a function of the peak intensity with electron refluxing taken into account at the foil thickness  $d = 5$  (1), 30 (2), 60 (3), 100 (4), 200 (5), 360 (6), and 3000  $\mu\text{m}$  (7) and  $\alpha_0 = 45^\circ$ .



**Figure 9.** Ratio of the  $K_\alpha$  yields from the front side of a thick (3 mm) and thin (10  $\mu\text{m}$ ) foils without electron refluxing as a function of the peak intensity at  $\alpha_0 = 45^\circ$ .



**Figure 7.**  $K_\alpha$  yield from the front (1, 3) and rear (2, 4) sides of the foil as a function of its thickness with electron refluxing taken into account at the peak intensity  $I_0 = 1.5 \times 10^{19}$  (1, 2) and  $10^{18}$  (3, 4)  $\text{W cm}^{-2}$  and  $\alpha_0 = \beta_0 = 45^\circ$ .

and absorption. This allows the calculation of the anisotropic yield of this emission from the front and back sides of the foil of arbitrary thickness with and without electron refluxing taken into account. We have proposed a method for diagnosing the effect of electron refluxing in the foils, which significantly increases the  $K_\alpha$  yield. The calculations and optimisation of the  $K_\alpha$  yield allow one to predict and analyse the results of measurements performed on the petawatt laser facilities. The developed model can also be used to refine the results of calculations of the characteristic X-ray yield from the target with a clustered surface [16] and compare them with the results of measurements [17].

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